

LUX-ZEPLIN Status



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XIV International Conference on Topics in Astroparticle and Underground Physics



LZ = LUX + ZEPLIN

32 institutions currently About 190 people

LIP Coimbra (Portugal)

MEPhI (Russia)

Edinburgh University (UK)

University of Liverpool (UK)

Imperial College London (UK)

University College London (UK)

University of Oxford (UK)

STFC Rutherford Appleton Laboratories (UK)

Shanghai Jiao Tong University (China)

University of Sheffield (UK)

University of Alabama

University at Albany SUNY

Berkeley Lab (LBNL)

University of California, Berkeley

Brookhaven National Laboratory

Brown University

University of California, Davis

Fermi National Accelerator Laboratory

Kavli Institute for Particle Astrophysics & Cosmology

Lawrence Livermore National Laboratory

University of Maryland

University of Michigan

Northwestern University

University of Rochester

University of California, Santa Barbara

University of South Dakota

South Dakota School of Mines & Technology

South Dakota Science and Technology Authority

SLAC National Accelerator Laboratory

Texas A&M

Washington University

University of Wisconsin

Yale University



Sanford Underground Research Facility



Davis Cavern 1480 m
(4200 mwe)
LZ in LUX Water Tank
South Dakota USA



LUX removed by early 2017 Water tank kept



Scale Up ≈ 50 in Fiducial Mass

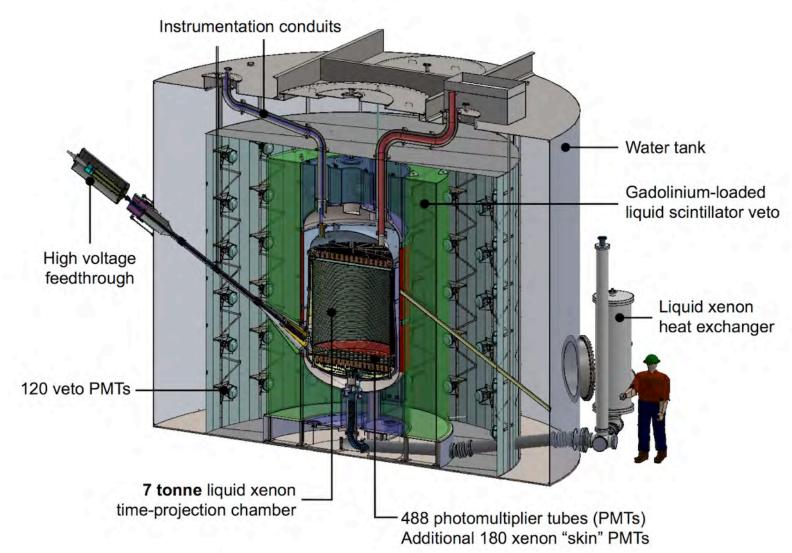
LZ

Total mass – 10 T
WIMP Active Mass – 7 T
WIMP Fiducial Mass – 5.6 T





LZ Overview





Key Design Points

- → 7 active tonnes of LXe can yield 2 x 10⁻⁴⁸ cm² sensitivity in about three years of running
- **★** 5.6 tonne fiducial volume, 1000 days
- ★ Requires all detector systems working together
 - □ Xe detector with good light collection, reasonable background rejection (ER discrimination) and good signal detection efficiency
 - □ Sophisticated veto system: skin (outside active Xe region)
 + scintillator/water allows maximum fiducial volume to be obtained, maximizes use of Xe and substantially increases reliability of background measurements
 - ☐ Control backgrounds, both internal (within the Xe) and external from detector components/environment

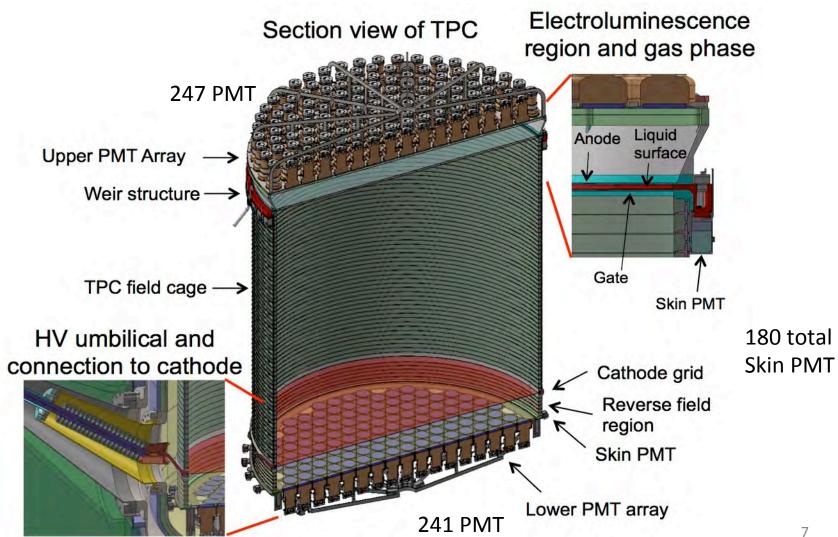


Design Status Summary

- ★ Conceptual, and in some cases more advanced design, completed for all aspects of detector
- ◆Conceptual Design Report about to appear on arXiv
- ★Acquisition of Xenon started
- → Procurement of PMTs and cryostat started
- ★Collaboration wide prototype program underway to guide and validate design
- → Backgrounds modeling and validation well underway



Xe TPC Detector





Xe Detector PMTs

- ★R11410-22 3" PMTs for TPC region
 - □ Extensive development program, 50 tubes in hand, benefit from similar development for XENON, PANDA-X and RED
 - ☐ Materials ordered and radioassays started prior to fabrication.
 - ☐ First production tubes early 2016.
 - □ Joint US and UK effort
- **★**R8520-406 1" for skin region
 - ☐ Considering using 2" or 3" for bottom dome region, recycle tubes from older detectors



Xe Detector Prototyping

- ★Extensive program of prototype development underway
- **→**Three general approaches
 - ☐ Testing in liquid argon, primarily of HV elements, at Yale and soon at LBNL
 - ☐ Design choice and validation in small (few kg) LXe test chambers in many locations: LLNL, Yale-> UC Berkeley, LBNL, U Michigan, UC Davis, Imperial College, MEPhI
 - ☐ System test platform at SLAC, Phase I about 100 kg of LXe, TPC prototype testing to begin in few months

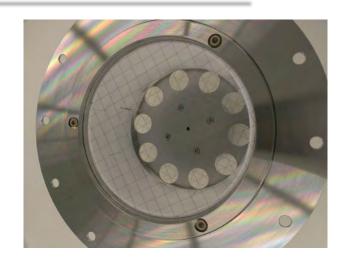


High Voltage Studies









Prototype of highest E-field region tested in LAr

- ◆ Cathode voltage design goal: 200 kV (provides margin)
- ◆ LZ nominal operating goal: 100 kV (~ 700 V/cm)
- ◆ Feedthrough prototype tested to 200 kV
- Prototype TPC for 100 kg LXe system fabrication starting
- ♦ HV prototyping expanding at Berkeley
- ◆ Wire electron emission studies at Imperial (see talk by A. Tomas)



Extensive Calibration

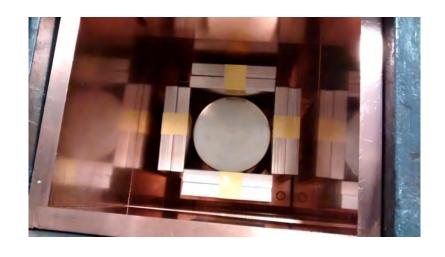
★LUX has led the way to detailed calibrations. LZ will build on this and do more.

Done in LUX and will be done in LZ	Not done in LUX, but will do in LZ
^{83m} Kr (routine, roughly weekly)	Activated Xe (129mXe and 131mXe)
Tritiated methane (every few months)	²²⁰ Rn
External radioisotope neutron sources	AmLi
External radioisotope gamma sources	YBe
DD neutron generator(upgraded early next year to shorten pulse)	



Cryostat Vessels

- → UK responsibility
- ★ Low background Ti chosen as cryostat material SS alternative advanced as backup
- ★ Ti slab for all vessels (and other parts) received and assayed
- → Contributes < 0.05 NR+ER counts in fiducial volume in 1,000 days after cuts
 </p>





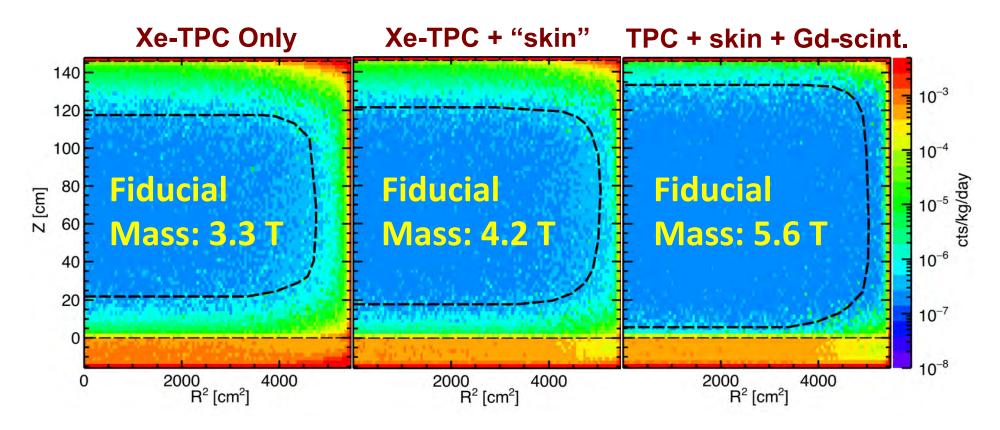




How to maximize the WIMP target mass?

- → Two-component outer detector:
 - □ 0.75 m thick Gd-loaded scintillator
 - □ instrumented Xenon "skin"
 - □ tag neutrons and gammas

in-situ monitoring of residual backgrounds!!!





LZ Projected Backgrounds

Expected backgrounds for 5.6 T fiducial - 1,000 days

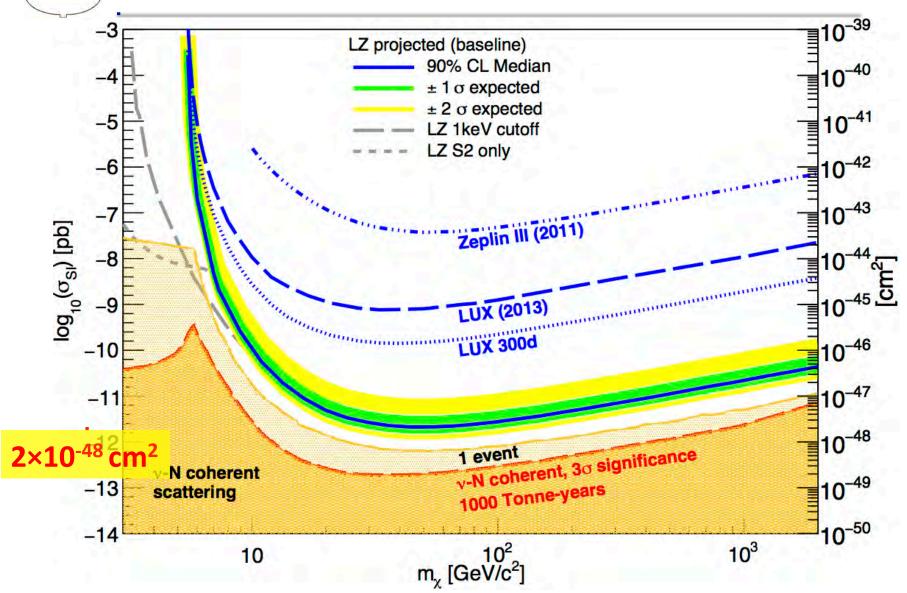
ER NR

Itom	Mass	U mBq/kg	Th mBq/kg	⁶⁰ Co mBq/kg	⁴⁰ K mBq/kg	n/yr	ER	NR
Item	kg						cts	cts
R11410 PMTs	93.7	2.7	2.0	3.9	62.1	373	1.24	0.20
R11410 bases	2.7	74.6	29.1	3.6	109.2	77	0.17	0.03
Cryostat vessels	2,140	0.09 1,507	0.23 1,065	≈0 ≈0 -	0.54 3,900 -	213 20,850 602	0.86 0.08 9.5	0.02 0.02 0.05
OD PMTs	122							
Other components								
Total components			1	144 4 9	10 _ 2 !!		11.9	0.32
Dispersed radionuclides	(Rn, Kr, Ar)						54.8	
¹³⁶ Xe 2νββ							53.8	10,70
Neutrinos (v-e, v-A)	4 1 10		4 - 41				271	0.5
Total events							391.5	0.82
WIMP background event		ceptance)		7			1.96	0.41
Total ER+NR background events					2.3	37		



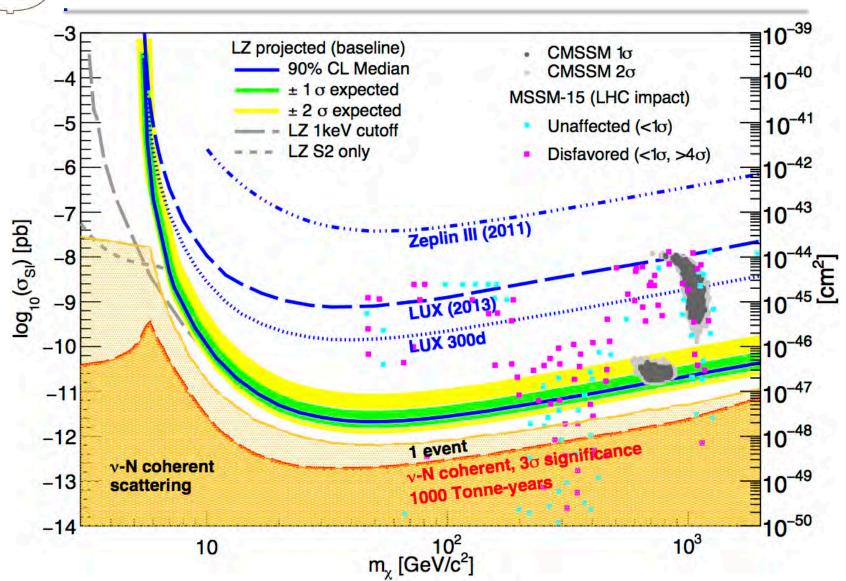
Projected Sensitivity – Spin Independent

(LZ 5.6 Tonnes, 1000 live days, 6 keV_{nr} analysis threshold)





Sensitivity with SUSY Theories





Neutrino Physics with LZ

+		Use self-shielding to reduce gamma-ray backgrounds in a 1-2 tonne fiducial mass Projected sensitivity: 90% confidence level $T^{0v}_{1/2}$ of 2 x 10^{26} years Enriching the Xe target could increase this to ~ 2 x 10^{27} years Current limit is 2.6 x 10^{25} years (preliminary) from KamLAND-Zen
+	Ext	ernal Neutrino Physics
		Solar neutrinos
		O Expect about 850 pp neutrino events between 1.5 and 20 keV _{ee}
		Supernova neutrinos
		O Via flavor-blind coherent neutrino-nucleus scattering
		O For a 10 kpc SN, LZ would see about 50 events with energy > 6 keV $_{\rm nr}$ and 100 events > 3 keV $_{\rm r}$
		Sterile neutrinos
		O Could use a 5 MCi ⁵¹ Cr source near LZ
		 Excellent position reconstruction for better source normalization, higher sterile neutrino masses.
		Neutrino magnetic moment

O Sensitivity near astrophysical limit of 2 x 10⁻¹² Bohr magnetons.



LZ Timeline

Year	Month	Activity
2012	March	LZ (LUX-ZEPLIN) collaboration formed
	May	First Collaboration Meeting
	September	DOE CD-0 for G2 dark matter experiments
2013	November	LZ R&D report submitted
2014	July	LZ Project selected in US and UK
2015	April	DOE CD-1/3a approval, similar in UK
		Begin long-lead procurements(Xe, PMT, cryostat)
2016	April	DOE CD-2/3b approval, baseline, all fab starts
2017	June	Begin preparations for surface assembly @ SURF
2018	July	Begin underground installation
2019	Feb	Begin commissioning
		18



Conclusions

- ★ LZ Project well underway, with procurement of Xe, PMTs and cryostat vessels started
- **★** Commissioning to begin in 2019
- **★** Extensive prototype program underway
- ★ LZ benefits from the excellent LUX calibration techniques and understanding of background
- ★ LZ sensitivity expected to reach 2×10⁻⁴⁸ cm²

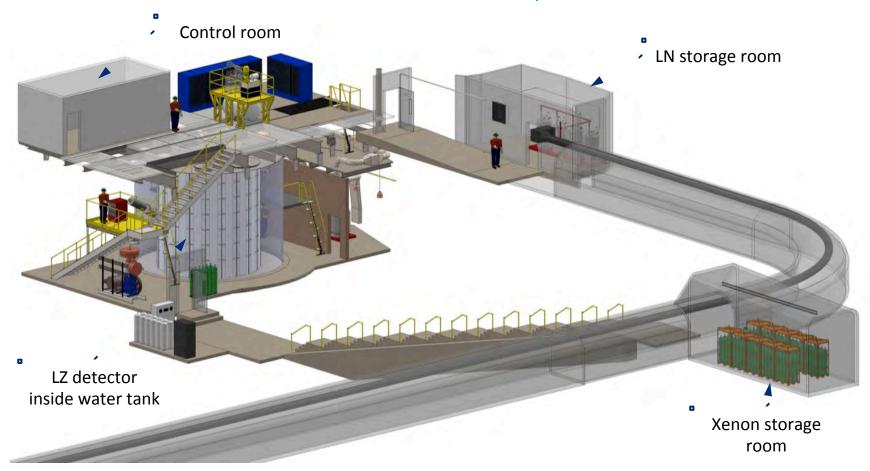


Extra Slides



LZ Underground at SURF

Years of experience at SURF from LUX





⁸⁵Kr Removal and Screening

- ightharpoonup Remove Kr to <15 ppq (10⁻¹⁵ g/g) using gas chromatography.
- → Best LUX batch 200 ppq
- **→** Setting up to process 200 kg/day at SLAC

→ Have a sampling program to instantly assay the removal at SLAC and

continuously assay in situ

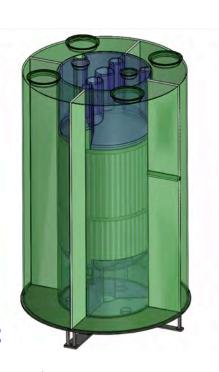


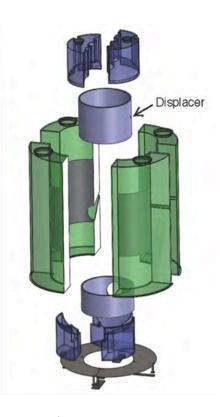




Outer Detector

- ★ Essential to utilize most Xe, maximize fiducial volume
- ★ Segmented tanks installation constraints (shaft, water tank)
- ★Gadolinium loaded scintillator, LAB, OK underground
- → Daya Bay legacy, scintillator & tanks (and people)
- Advanced conceptual design



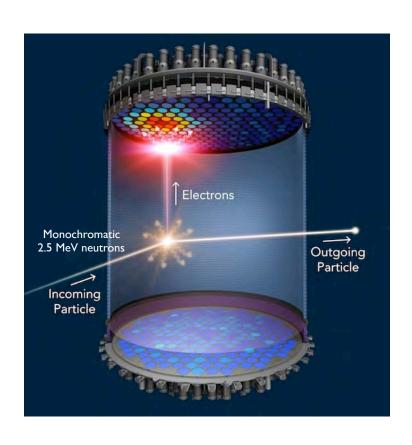


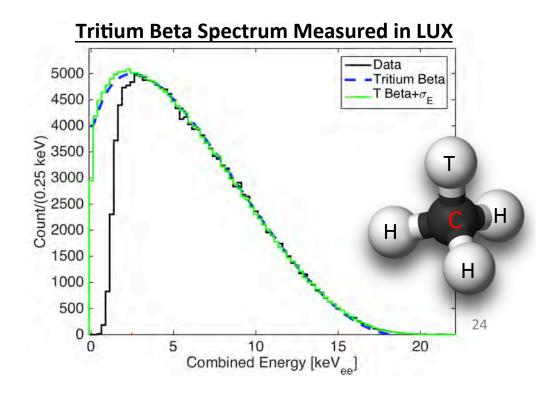
Layout of the LZ outer detector system, which consists of nine acrylic tanks. The largest are the four quarter-tanks on the sides. Two tanks cover the top, and three the bottom. The exploded view on the right shows the displacer cylinders placed between the acrylic vessels and the cryostat.



LZ Calibrations

- → Demonstrated in LUX. Calibrate The Signal and Background Model in situ.
- **→** DD Neutron Generator (Nuclear Recoils)
- **★**Tritiated Methane (Electron Recoils)
- **★**Additional Sources e.g. YBe Source for low energy (Nuclear Recoils)

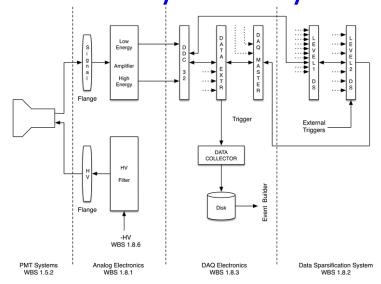




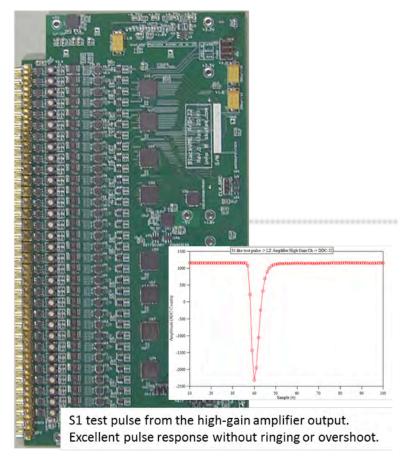


Electronics/DAQ

- ★LUX legacy, augmented by experienced new groups (primarily DAQ, controls)
- ◆ Prototyping underway, will lead to full – chain test of key elements by end of year



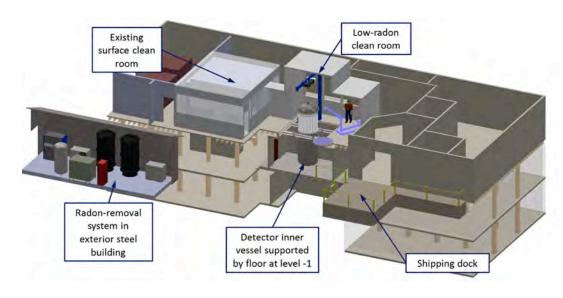
32 channel digitizer prototype





Integration/Installation

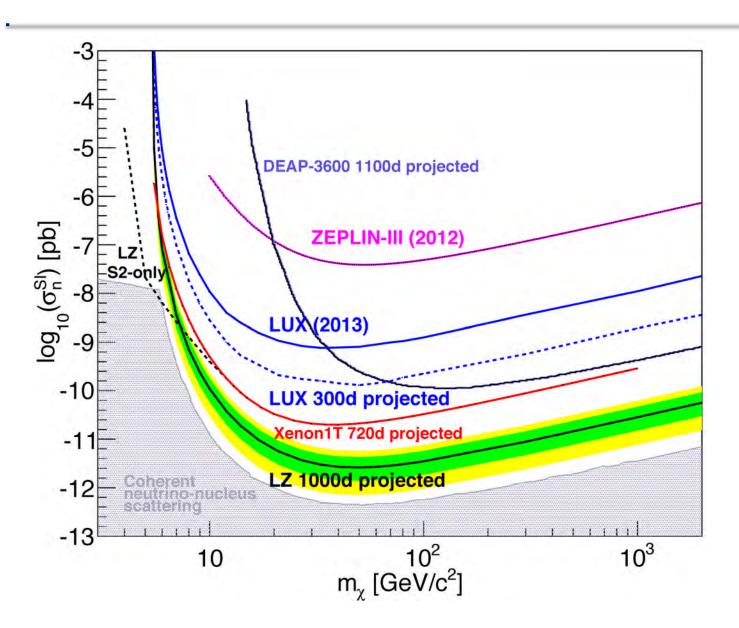
- ★Surface assembly of TPC into inner cryostat
- **★**LUX experience at SURF
- → Dedicated on site infrastructure improvements for LZ. Design started, construction





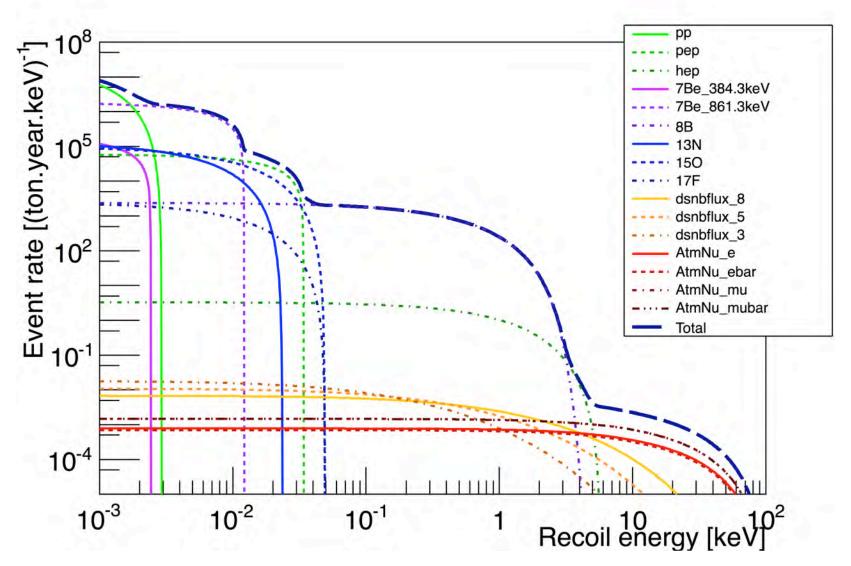


Sensitivity with Competition



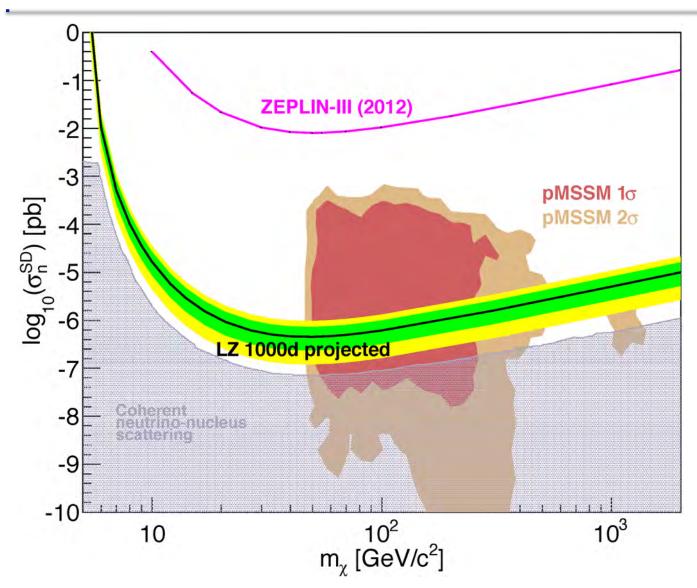


Response of Xe to Neutrinos arXiv:1307:5458



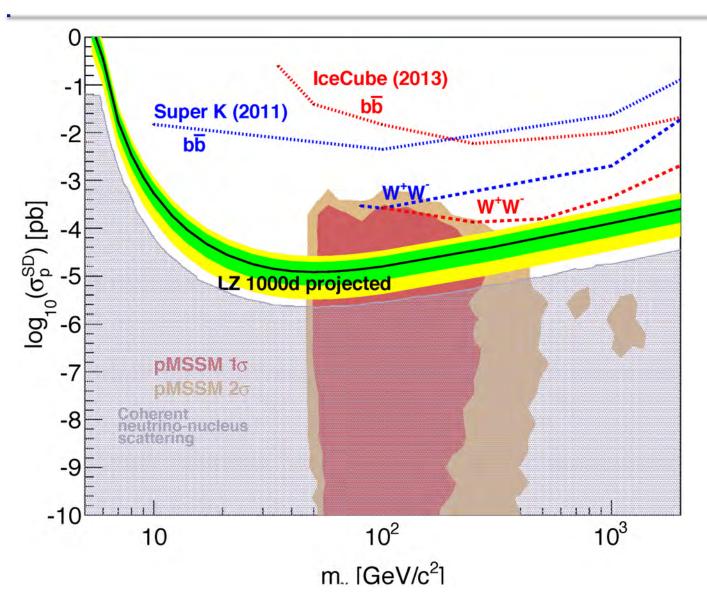


Spin Dependent Neutron

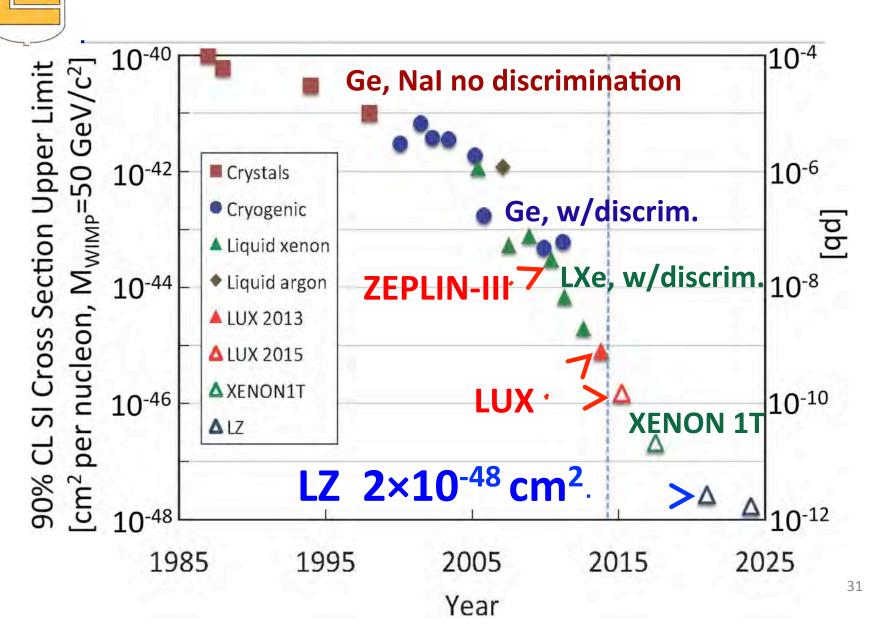




Spin Dependent Proton



Time Evolution





Running Time

- ★Sensitivity vs. running time.
- ◆1,000 days is the nominal.
- **→** Baseline backgrounds
- ★Rapid improvement in sensitivity
- ◆ Potential to eventually get to ~ 1 x 10⁻⁴⁸ cm²

