

#### PHYSICS REACH OF THE LZ EXPERIMENT Amy Cottle, University of Oxford TAUP 2021





# Introduction to LUX-ZEPLIN (LZ)



- \* LZ is a dark matter direct detection experiment based at the Sanford Underground Research Facility
  - 4850 ft underground -> cosmic ray reduction
- \* Dual-phase xenon time projection chamber (TPC)
  - 7 t active xenon viewed by 494 photomultiplier tubes
  - Skin and Outer Detector (OD) active veto systems
- \* Commissioning started (see <u>D. Woodward's talk</u>)
- Primarily designed for WIMP detection, but has considerable sensitivity to other new physics

<u>j.nima.2019.163047</u>



## **TPC Detection Principle**

- \* Interactions in the xenon create
  - Light prompt scintillation (S1)
  - Charge electrons drifted and extracted into gas -> proportional scintillation (S2)
- Excellent 3D position reconstruction (~mm)
  - Z from time difference between S1 and S2
  - XY from S2 hit pattern on top PMT array
- Signals can be used to determine energy of the event and for particle identification



# WIMP Analysis Strategy



Simulated data inside the fiducial volume for the full LZ exposure (1000 days \* 5.6 t)

- S2 to S1 ratio -> discrimination of electron recoil (ER) from nuclear recoil (NR) events
  - ER and NR bands obtained via calibration
  - Conduct search below NR band median -> predicted 99.5% ER event rejection
- \* Reduce background events via
  - Fiducialisation to cut external source contributions
  - Rejection if coincident veto response (<u>Slide 7</u>)
- \* Projected energy threshold of ~1.5 keVee



# Background Control

- \* Material selection based on ~2000 assays with 13 HPGe detectors, ICPMS, neutron activation analysis
- Four Rn emanation screening sites
- \* TPC assembly in Rn-reduced cleanroom
  - Cleanliness protocols limiting surface contamination
    - Dust ( $< 500 \text{ ng/cm}^3$ )





# **Background Simulations**

- \* Geant4-based simulation framework (BACCARAT)
  - Event generators written for specific backgrounds e.g.
    - Neutrons with coincident gammas (e.g. uranium spontaneous fission)
      Laboratory and external backgrounds (e.g. muons, covern gammas)
    - Laboratory and external backgrounds (e.g. muons, cavern gammas)
    - Surface backgrounds (e.g. embedded Po210 (alpha, n) on fluorine)
  - Custom physics lists e.g. OD Gd neutron capture with DICEBOX
- \* Energy depositions converted to S1s and S2s with NEST
  - Used for backgrounds estimates and sensitivity projections
- \* Detailed optical and electronics response simulations also possible to produce pulses and realistic scatter topologies for mock analyses









CAD view of the Outer Detector

## WIMP Search Backgrounds



Expected contributions after analysis cuts (single scatter (SS), 5.6 t fiducial volume (FV), veto anti-coincidence)



#### Backgrounds for a 40 GeV/c<sup>2</sup> WIMP



Total

+ 99.5% ER discrimination, 50% NR efficiency

ER [cts]	NR [cts]
1131	1.03
5.66	0.52





## WIMP Sensitivity

#### Profile likelihood ratio analysis S1, S2 (+ position) PDFs Simulated backgrounds (<u>Slide 6</u>) Analytical recoil spectra -> NEST (signal/physics) 90% CL minimum: $1.4 \times 10^{-48} \text{ cm}^2 \text{ at } 40 \text{ GeV}/c^2$

For full LZ exposure (1000 days \* 5.6 t)

PRD 101, 052002 (2020)





#### Extension to Lower Mass Candidates

- \* Lowering the energy threshold
  - Exploit 2-phe effect to reduce S1 coincidence requirement from 3 to 2

~4x sensitivity gain at  $2.5 \text{ GeV}/c^2$ •

- Conduct an S2-only search
  - Discriminate backgrounds based on • S2 pulse shape/width
- Sub-GeV masses accessible when considering Migdal electron emission

ArXiv: 2101.08753





Image credit: <u>PRL 121, 101801</u>

# Low Energy ER Searches

- Signal & background models use NEST reconstructed energy, S1, S2
- \* Example 90% CL upper limits:
  - Solar axions  $g_{Ae} < 1.58 \times 10^{-12}$
  - Neutrino magnetic moment  $\mu_{vsolar} < 6.2 \times 10^{-12} \mu_{B}$
  - Neutrino electric millicharge  $q_{v_{solar}} < 1.4 \times 10^{-13} e_0$
- \* <sup>37</sup>Ar & <sup>3</sup>H included in likelihood fits
- Robustly test XENON1T excess
  <u>ArXiv: 2102.11740</u>



#### 136Xe Neutrinoless Double Beta Decay

- \* Isotopic abundance of 8.9% <sup>136</sup>Xe
- \* Gamma backgrounds near  $Q_{\beta\beta}$  (2458 keV)
- Energy and position dependent PDFs
- \* Requires good energy resolution (assume 1% or better at  $Q_{\beta\beta}$ )
- \* Needs good single vs multiple scatter discrimination (Z separation < 3 mm)

 $T_{1/2}$  (90% C.L.) > 1 x 10<sup>26</sup> years in 1000 live days in 5.6 t fiducial volume Rate [counts/kg/day/keV]

PRC 102, 014602 (2020)



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#### Conclusions



- \* LZ is fully optimised for the direct detection of WIMPs
  - 20x more sensitive than the current best limits
- \* LZ is a multi-physics experiment, with competitive searches across a range of energies, in both NR & ER channels
- \* Long-term campaign of backgrounds control and continued assessment to ensure world-leading sensitivities
- Experiment is in its commissioning phase, with first science data expected this year
- This next year will be a pivotal one in dark matter physics watch this space!



#### LZ (LUX-ZEPLIN) Collaboration 34 Institutions: 250 scientists, engineers, and technical staff

- Black Hills State University
- Brandeis University
- Brookhaven National Laboratory
- Brown University
- Center for Underground Physics
- Edinburgh University
- Fermi National Accelerator Lab.
- Imperial College London
- Lawrence Berkeley National Lab.
- Lawrence Livermore National Lab.
- LIP Coimbra
- Northwestern University
- Pennsylvania State University
- Royal Holloway University of London
- SLAC National Accelerator Lab.
- South Dakota School of Mines & Tech
- South Dakota Science & Technology Authority
- STFC Rutherford Appleton Lab.
- Texas A&M University
- University of Albany, SUNY
- University of Alabama
- University of Bristol
- University College London
- University of California Berkeley
- University of California Davis
- University of California Santa Barbara
- University of Liverpool
- University of Maryland
- University of Massachusetts, Amherst
- University of Michigan
- University of Oxford
- University of Rochester
- University of Sheffield
- University of Wisconsin, Madison





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