



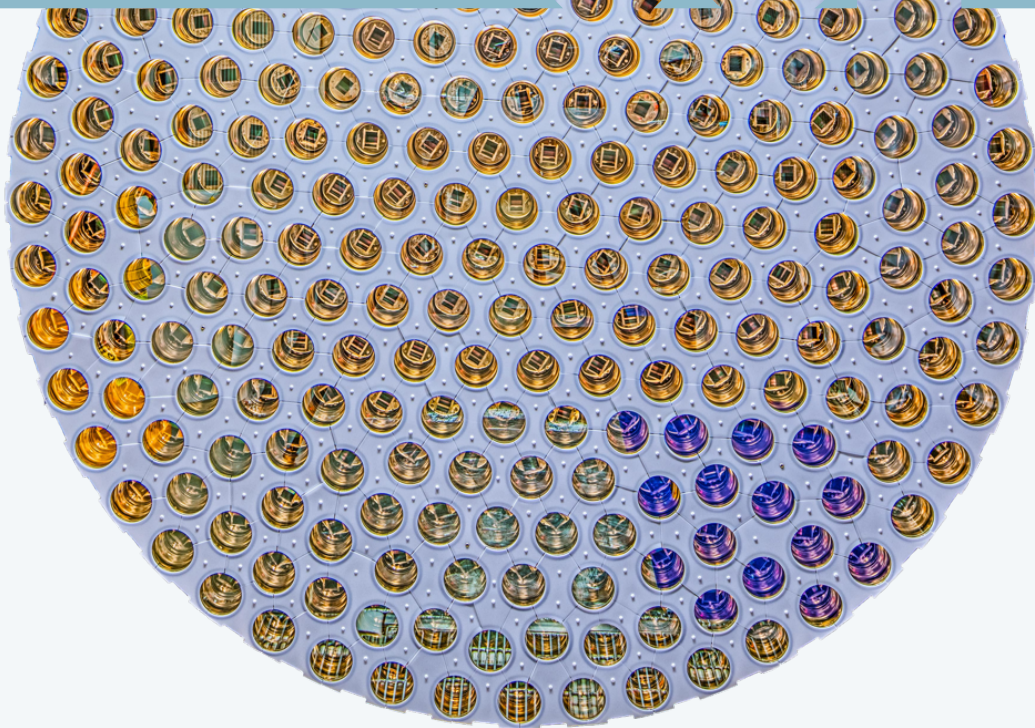
Status of the LZ Experiment

Theresa Fruth (UCL)

IoP 2021

Joint APP, HEP and NP conference

13th April 2021



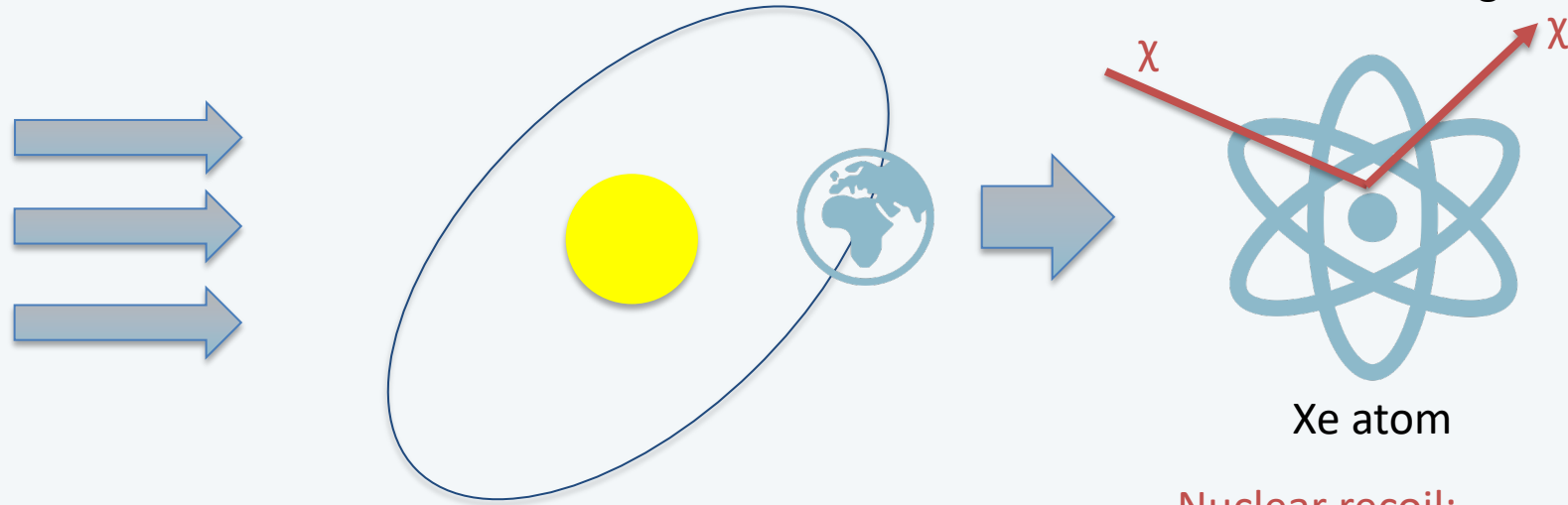
LZ collaboration

- 1) Black Hills State University
 - 2) Brandeis University
 - 3) Brookhaven National Laboratory
 - 4) Brown University
 - 5) Center for Underground Physics
 - 6) Edinburgh University
 - 7) Fermi National Accelerator Lab.
 - 8) Imperial College London
 - 9) Lawrence Berkeley National Lab.
 - 10) Lawrence Livermore National Lab.
 - 11) LIP Coimbra
 - 12) Northwestern University
 - 13) Pennsylvania State University
 - 14) Royal Holloway University of London
 - 15) SLAC National Accelerator Lab.
 - 16) South Dakota School of Mines & Tech
 - 17) South Dakota Science & Technology Authority
 - 18) STFC Rutherford Appleton Lab.
 - 19) Texas A&M University
 - 20) University of Albany, SUNY
 - 21) University of Alabama
 - 22) University of Bristol
 - 23) University College London
 - 24) University of California Berkeley
 - 25) University of California Davis
 - 26) University of California Santa Barbara
 - 27) University of Liverpool
 - 28) University of Maryland
 - 29) University of Massachusetts, Amherst
 - 30) University of Michigan
 - 31) University of Oxford
 - 32) University of Rochester
 - 33) University of Sheffield
 - 34) University of Wisconsin, Madison
- US UK Portugal Korea



Supported by DOE, STFC & SURF

Dark Matter wind



Elastic scattering:

Xe atom

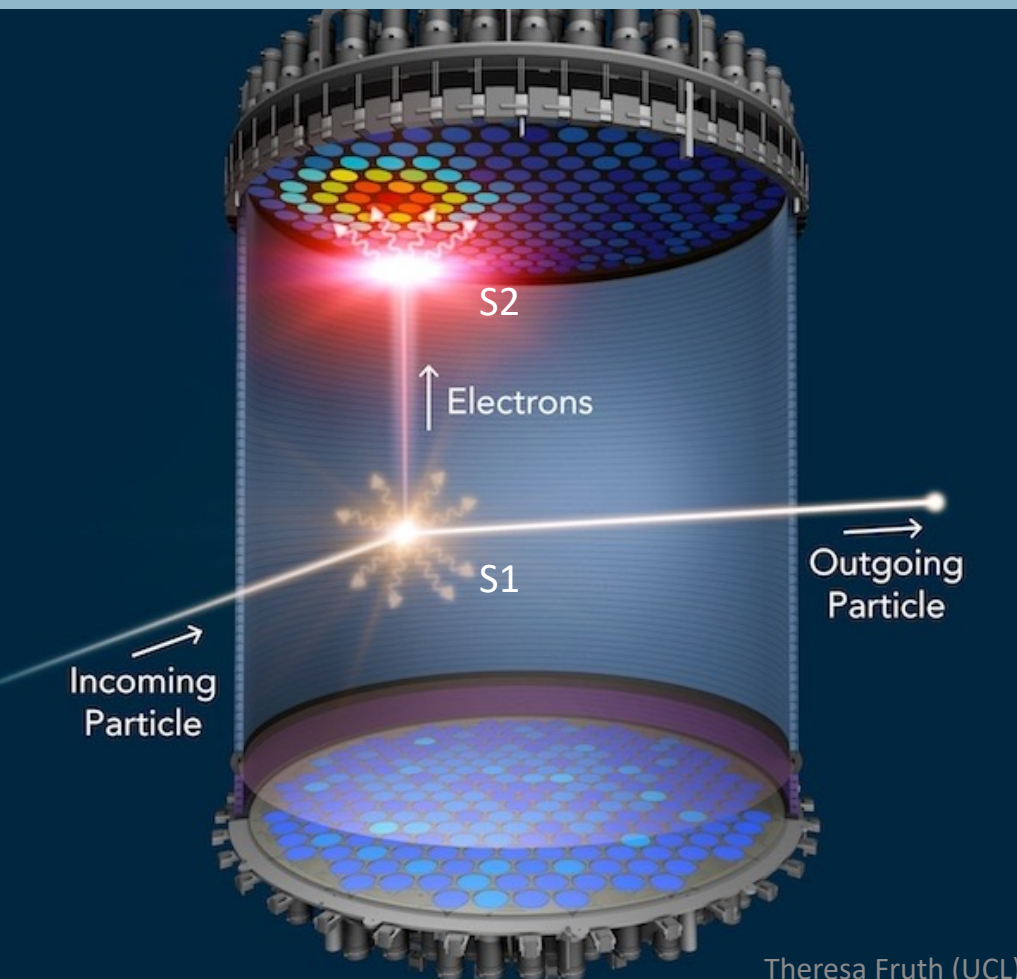
Milky Way – DM Halo:

- Density near sun $\sim 0.3 \text{ GeV/cm}^3$
- Mean particle speed $v \sim 300 \text{ km/s}$

Nuclear recoil:

- Ionization (charge)
- Scintillation (light)

TPC Overview

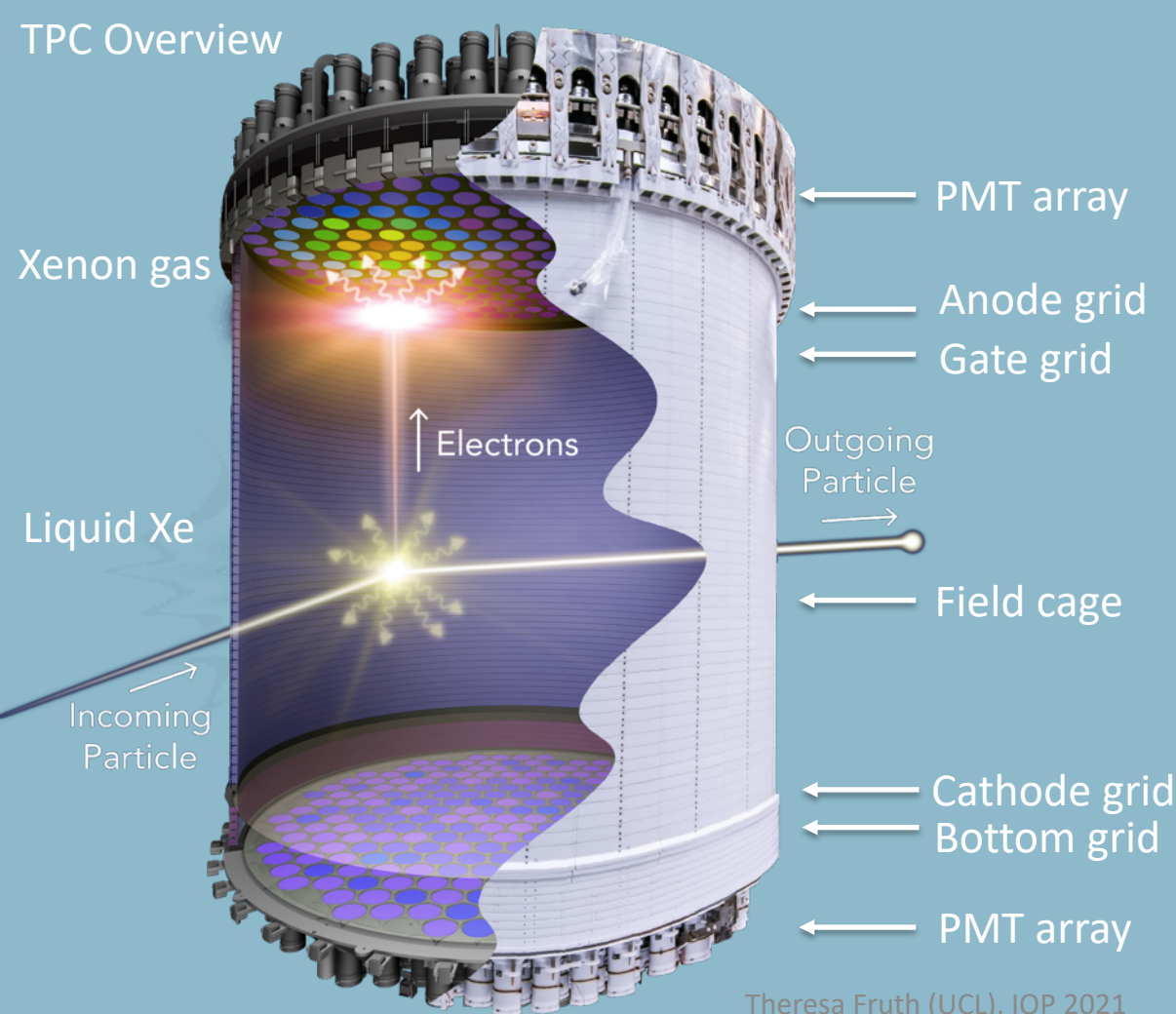


time

S2 – electroluminescence signal

Drift time
indicates depth

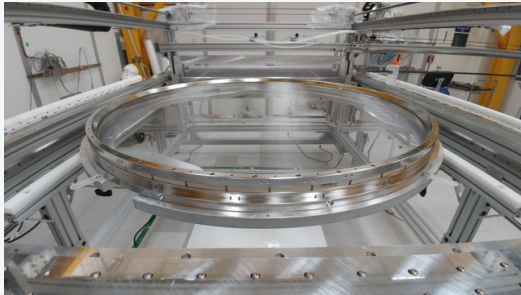
S1 – scintillation signal



- 7 tonnes of Xenon, 5.6 tonnes fiducial volume
- Interaction leads to prompt scintillation and free electrons
- Electric field to extract electrons into gas leading to Electroluminescence light
- 3D reconstruction with S2 (XY) and S1-S2 delay (Z) allows fiducialisation



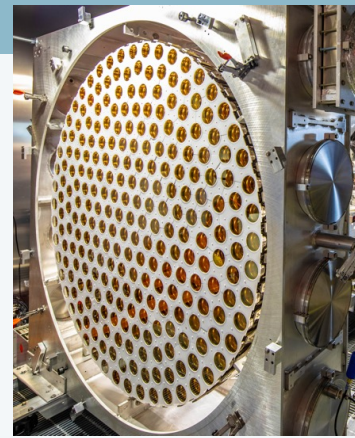
Bottom PMT array with field cage



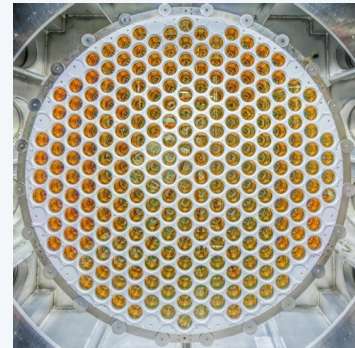
HV grid weaving at SLAC



Assembled TPC (July 2019)

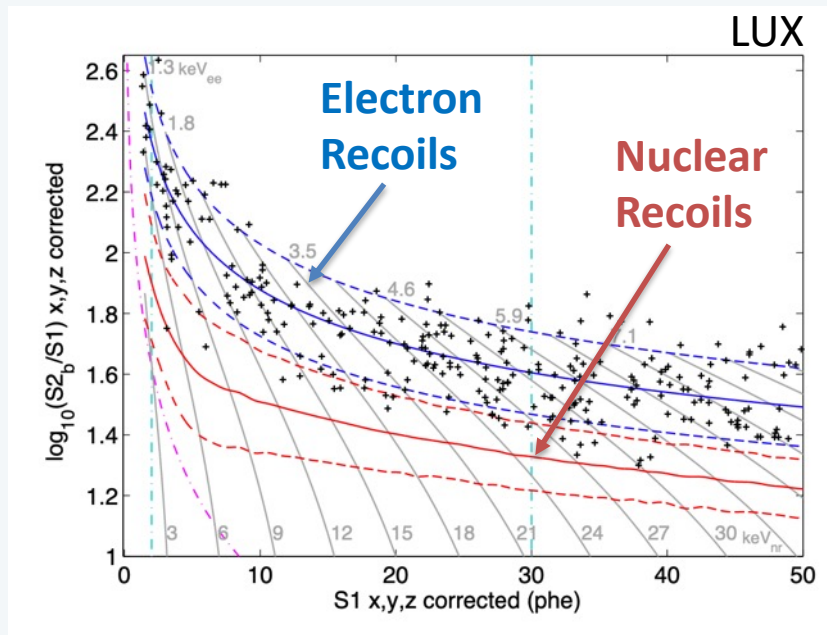


Top PMT array



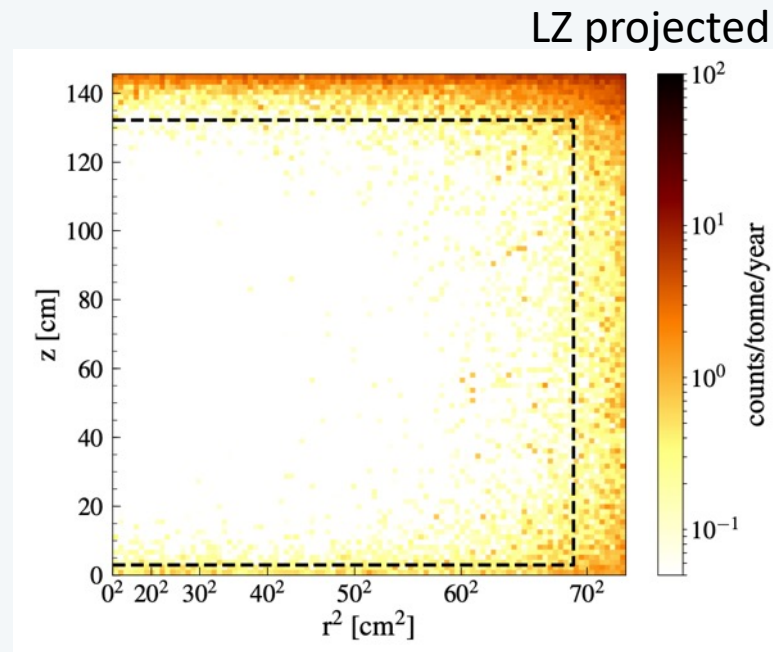
Bottom PMT array

ER-NR discrimination



[Phys. Rev. Lett. 112, 091303 \(2014\)](#)

Fiducialization



[Phys. Rev. D 101, 052002 \(2020\)](#)

Material Selection:

- Radio-assay campaign with gamma-screening and ICPMS
- Radon emanation:
 - 4 Rn emanation screening sites
 - Target Rn activity: 2 $\mu\text{Bq/kg}$

Xenon purification:

- Charcoal chromatography @ SLAC to remove Xenon contaminants – ^{85}Kr and ^{39}Ar
- Online gas purification at 500 slpm, turnover of total volume every 2.5 days

Cleanliness during construction:

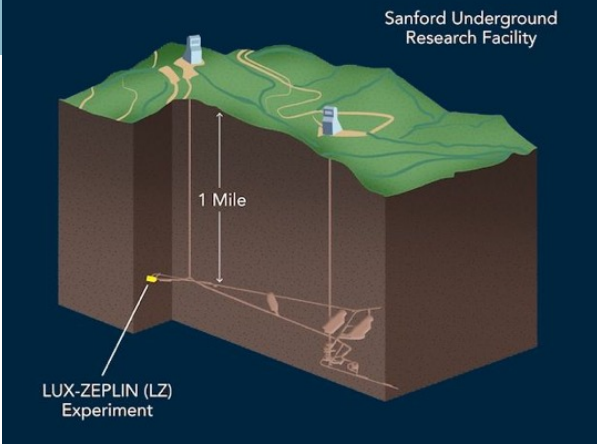
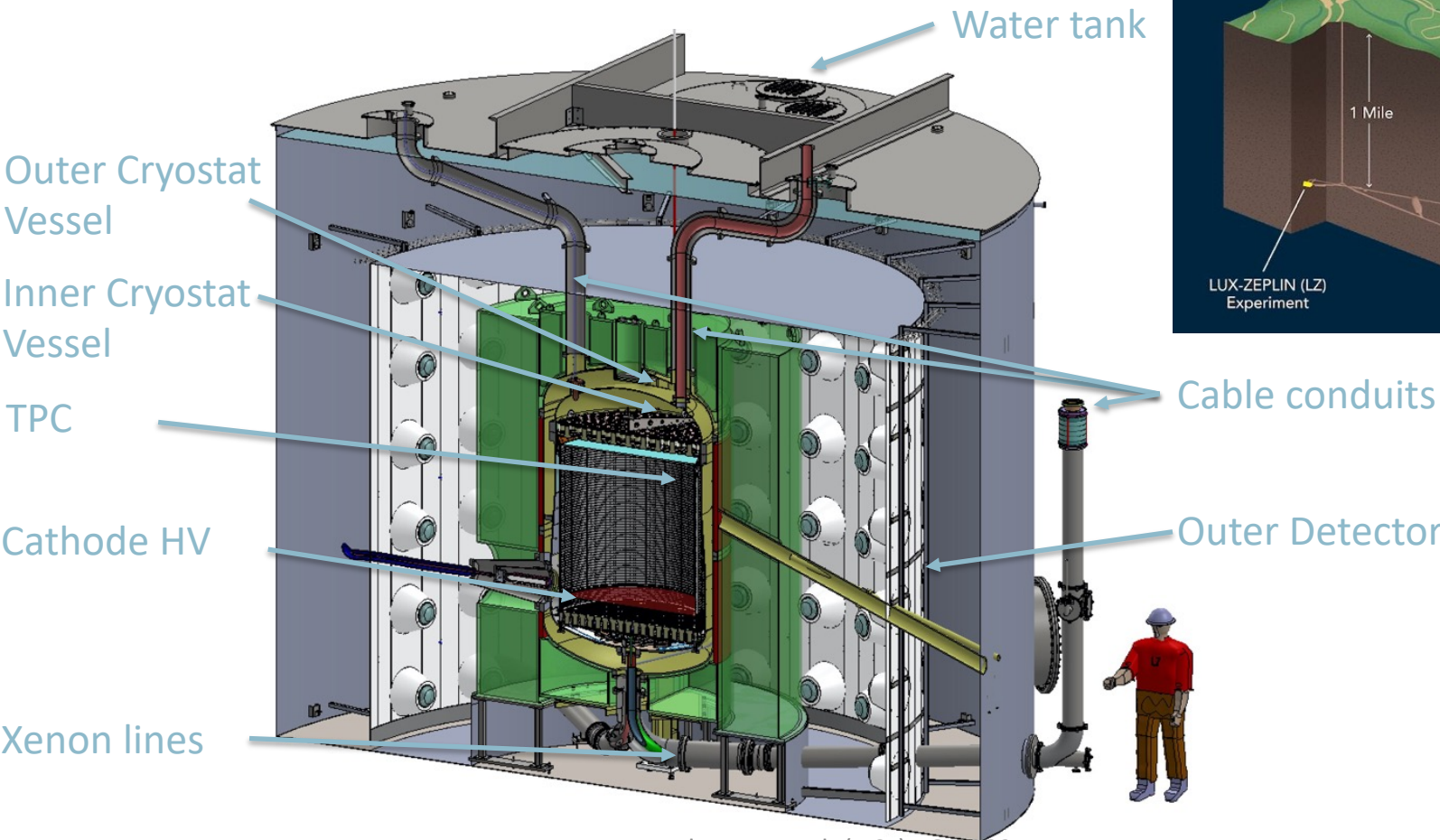
- Rn daughters and dust on surfaces
- TPC assembly in Rn-reduced cleanroom
- Dust $< 500 \text{ ng/cm}^2$ on all LXe wetted surfaces
- Rn-daughter plate-out on TPC walls $< 0.5 \text{ mBq/m}^2$

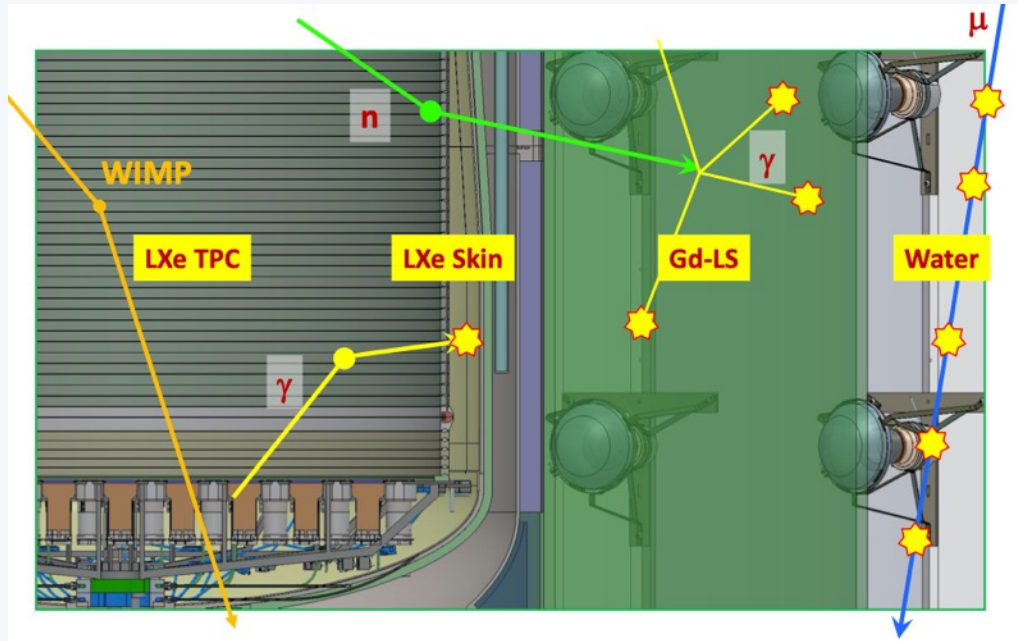
See Assembling the LZ detector
by N. Angelides (Tuesday, 3.45 pm)

Shielding:

- Deep underground
- High purity water shield
- Veto detectors

Integrated detector overview





Skin:

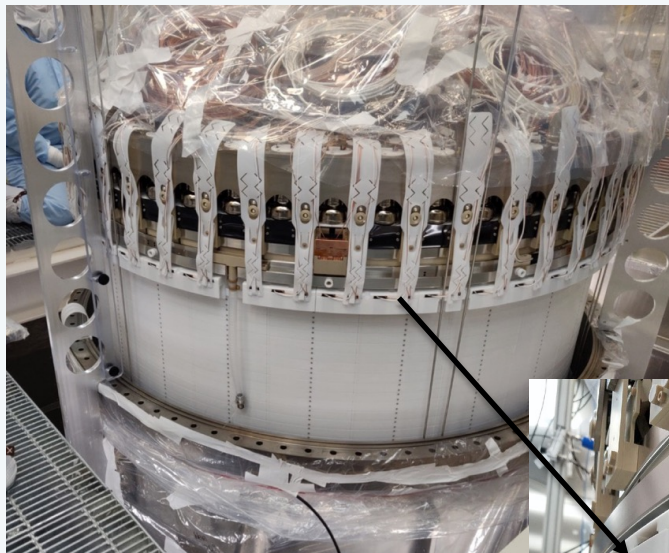
- 2 tonnes of LXe surrounding the TPC
- PMTs at top and bottom of the skin region
- Lined with PTFE to maximize light collection efficiency
- Anti-coincidence detector for γ -rays

Outer detector:

- 17 tonnes Gd-loaded liquid scintillator in acrylic vessels
- 120 8" PMTs mounted in the water tank
- Anti-coincidence detector for γ -rays and neutrons
- Observe ~ 8 MeV γ -rays from thermal neutron capture



PTFE tiled cryostat with bottom skin PMTs



TPC insertion into the Inner Cryostat, August 2019



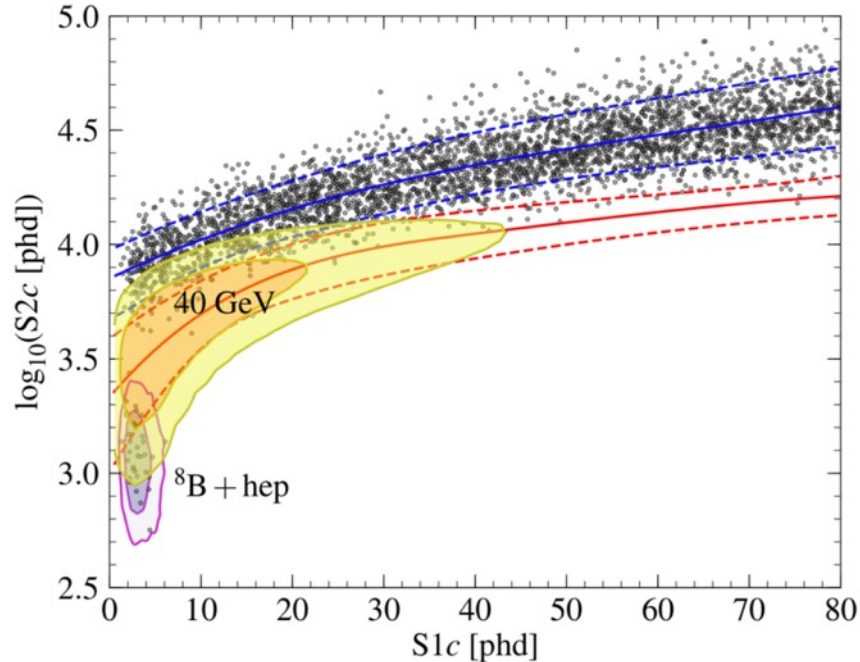
Top side skin PMTs



OD tanks and OCV in the water tank

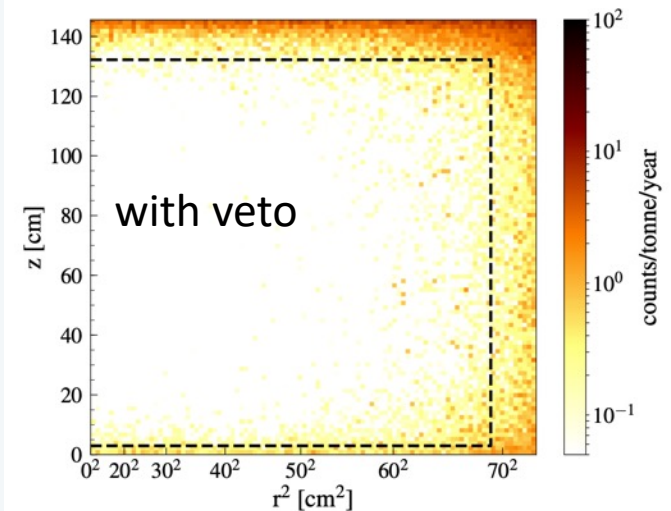
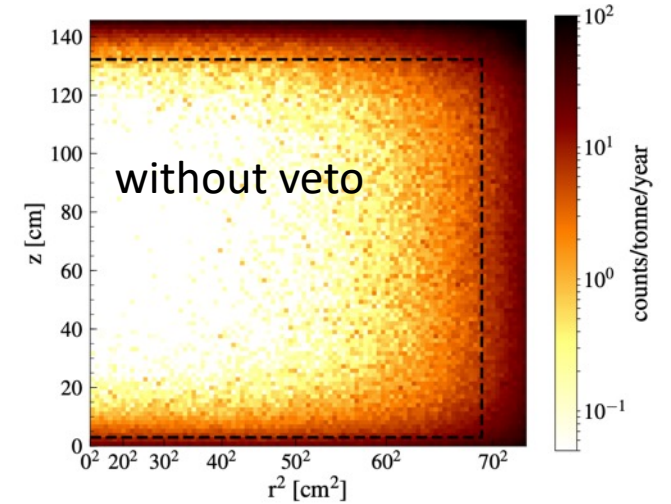
1000 day science run – expected backgrounds

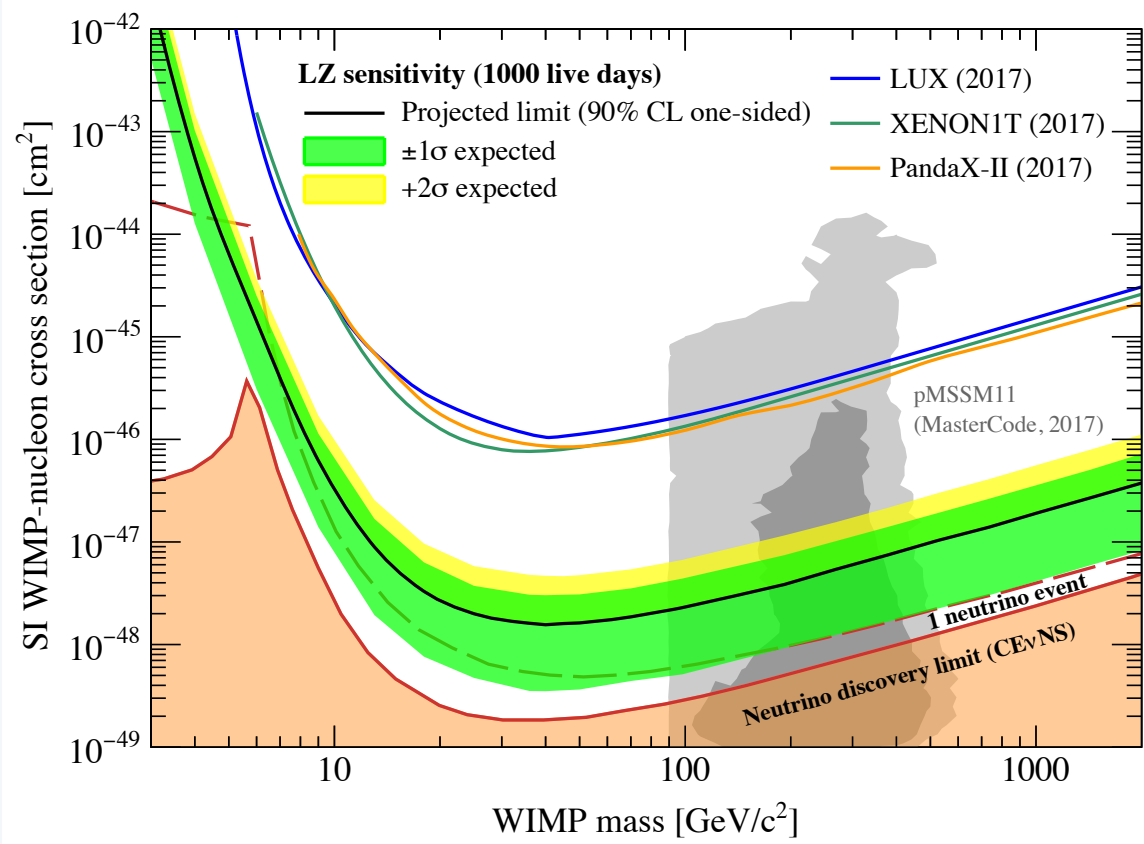
[Phys. Rev. D 101, 052002 \(2020\)](#)



1000 day run – after discrimination:
5.97 ER events and 0.51 NR events

Theresa Fruth (UCL), IOP 2021



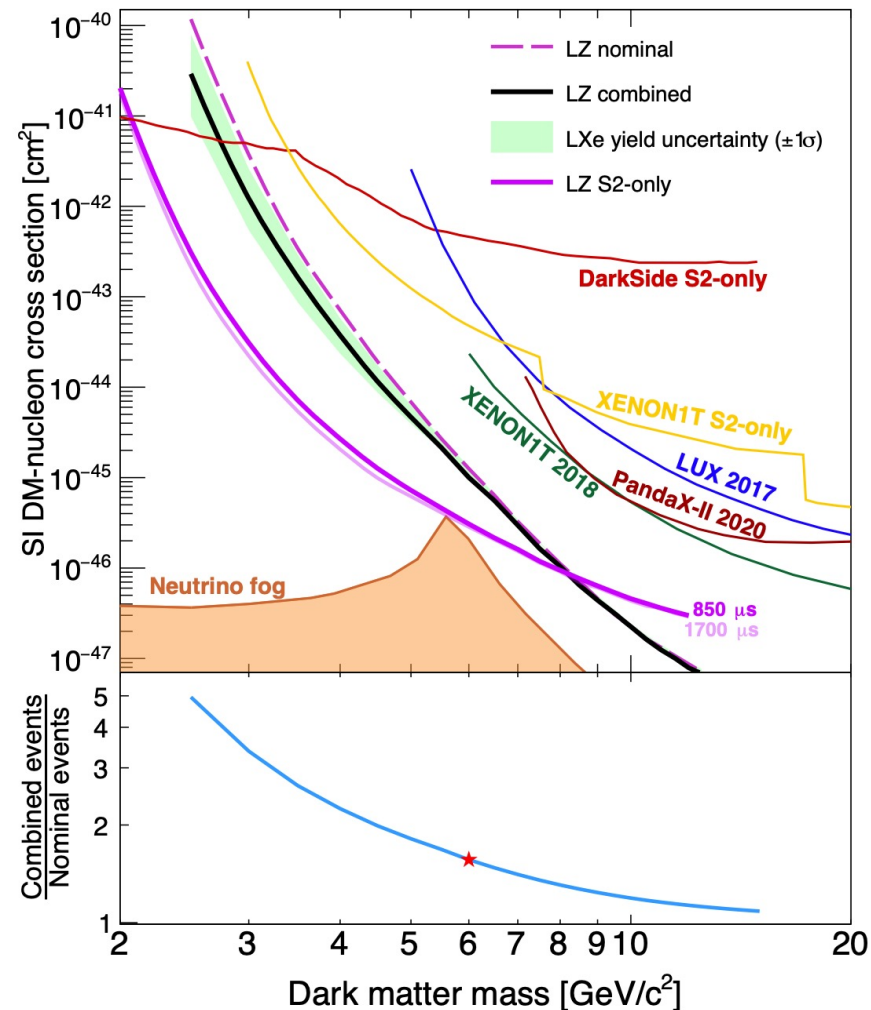


90% CL minimum of
 $1.4 \times 10^{-48} \text{ cm}^2$
at $40 \text{ GeV}/c^2$

Enhancing sensitivity in the low WIMP mass regime by lower in the detector threshold.

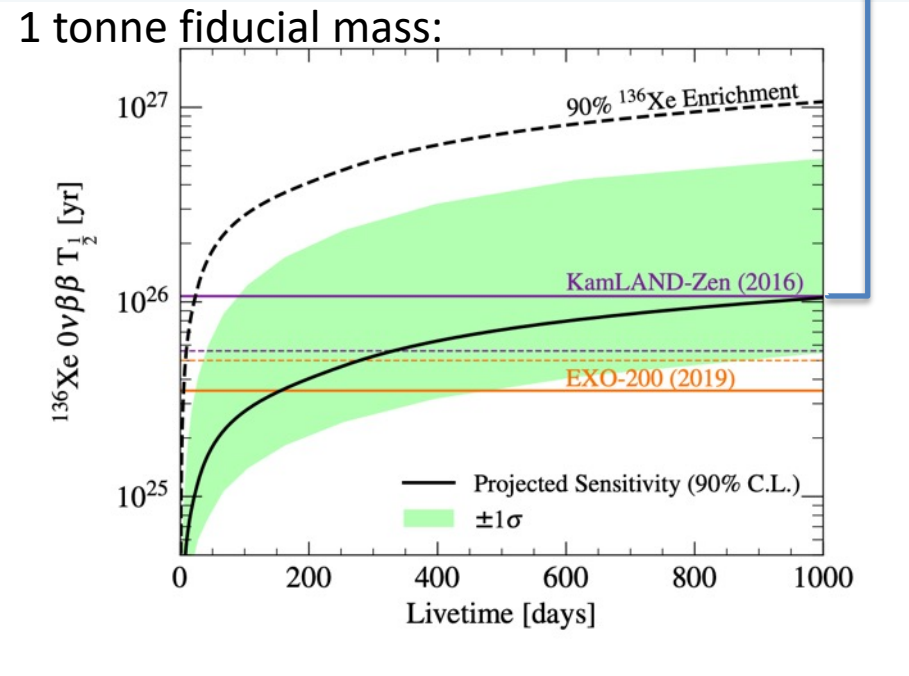
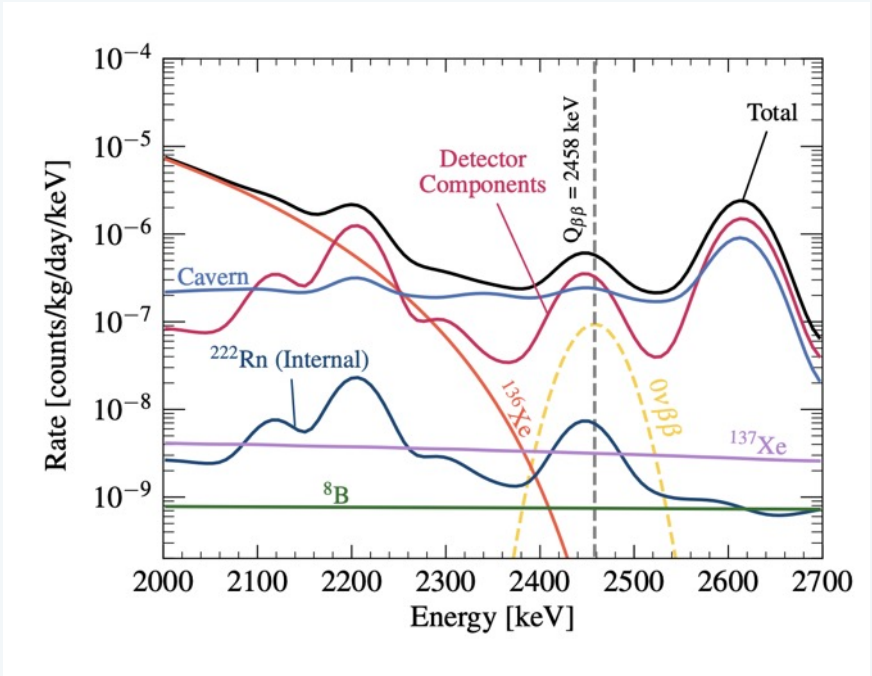
- Lower S1 coincidence requirements from 3 to 2 photons (making use of the DPE effect) -> LZ combined
- S2-only analysis -> LZ S2-only (for nominal and enhanced electron lifetime)

<https://arxiv.org/pdf/2101.08753.pdf>



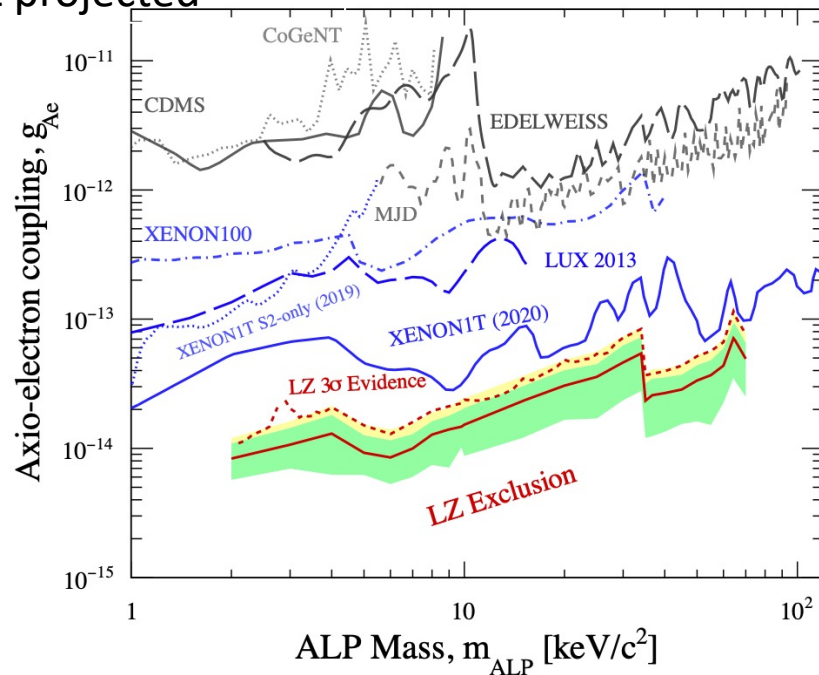
$^{136}\text{Xe } Q_{\beta\beta} = 2458 \text{ keV}$

$T_{1/2} \text{ (90\% C.L.)} > 1.06 \times 10^{26} \text{ years in 1000 live-days}$



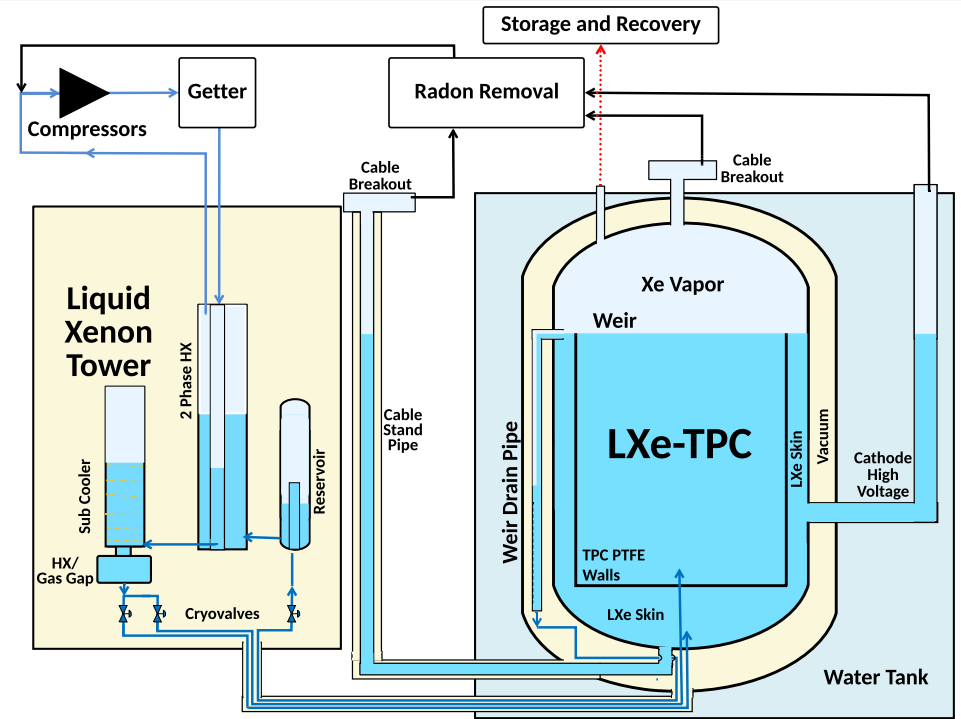
See E. Leason's talk about ER band searches with LZ (Tuesday, 3:00 PM)

LZ projected



- ER band searches for axions and ALPs assuming axio-electric interaction
- ALPs – monoenergetic feature in ER band
- Plot shows expected sensitivity for 1000 live-days and 5.6 tonne fiducial mass.





[NIM A953 \(2020\)163047](#)



Outlook

- Good progress in assembly and integration of detector and associated systems
- Expecting first data later this year
- Expected WIMP sensitivity of $1.4 \times 10^{-48} \text{ cm}^2$ at $40 \text{ GeV}/c^2$
- Also sensitive to a range of non-WIMP physics
- Stay tuned!

Acknowledgment of support:



Thank you!

Other LZ contributions this IoP conference:

Elizabeth Leason, New Physics Searches with Low Energy Electron Recoils in the LUX ZEPLIN Experiment (Tuesday, 3:00 PM, Astroparticle 2)

Nicolas Angelides, Assembling the LZ Detector (Tuesday, 3:45 PM Astroparticle 2)

Sam Eriksen, LUX-ZEPLIN optical photon simulations using GPUs (Thursday, 11:45 AM, Astroparticle 3)

Tom Rushton, Analysis of Neutron Activation Background in the LUX-ZEPLIN Experiment (Poster)

Expected backgrounds – 1000 day science run

TABLE III. Estimated backgrounds from all significant sources in the LZ 1000 day WIMP search exposure. Counts are for a region of interest relevant to a 40 GeV/c² WIMP: approximately 1.5–6.5 keV for ERs and 6–30 keV for NRs; and after application of the single scatter, skin and OD veto, and 5.6 tonne fiducial volume cuts. Mass-weighted average activities are shown for composite materials and the ²³⁸U and ²³²Th chains are split into contributions from early- and late-chain, with the latter defined as those coming from isotopes below and including ²²⁶Ra and ²²⁴Ra, respectively.

Background Source	Mass (kg)	²³⁸ U _e	²³⁸ U _l	²³² Th _e	²³² Th _l	⁶⁰ Co	⁴⁰ K	n/yr	ER (cts)	NR (cts)
		mBq/kg								
Detector Components										
PMT systems	308	31.2	5.20	2.32	2.29	1.46	18.6	248	2.82	0.027
TPC systems	373	3.28	1.01	0.84	0.76	2.58	7.80	79.9	4.33	0.022
Cryostat	2778	2.88	0.63	0.48	0.51	0.31	2.62	323	1.27	0.018
Outer detector (OD)	22950	6.13	4.74	3.78	3.71	0.33	13.8	8061	0.62	0.001
All else	358	3.61	1.25	0.55	0.65	1.31	2.64	39.1	0.11	0.003
subtotal									9	0.07
Surface Contamination										
Dust (intrinsic activity, 500 ng/cm ²)									0.2	0.05
Plate-out (PTFE panels, 50 nBq/cm ²)									-	0.05
²¹⁰ Bi mobility (0.1 μBq/kg LXe)									40.0	-
Ion misreconstruction (50 nBq/cm ²)									-	0.16
²¹⁰ Pb (in bulk PTFE, 10 mBq/kg PTFE)									-	0.12
subtotal									40	0.39
Xenon contaminants										
²²² Rn (1.8 μBq/kg)									681	-
²²⁰ Rn (0.09 μBq/kg)									111	-
^{nat} Kr (0.015 ppt g/g)									24.5	-
^{nat} Ar (0.45 ppb g/g)									2.5	-
subtotal									819	0
Laboratory and Cosmogenics										
Laboratory rock walls									4.6	0.00
Muon induced neutrons									-	0.06
Cosmogenic activation									0.2	-
subtotal									5	0.06
Physics										
¹³⁶ Xe 2νββ									67	-
Solar neutrinos: pp+ ⁷ Be+ ¹³ N, ⁸ B+hep									191	0*
Diffuse supernova neutrinos (DSN)									-	0.05
Atmospheric neutrinos (Atm)									-	0.46
subtotal									258	0.51
Total									1131	1.03
Total (with 99.5% ER discrimination, 50% NR efficiency)									5.66	0.52
Sum of ER and NR in LZ for 1000 days, 5.6 tonne FV, with all analysis cuts									6.18	

* Below the 6 keV NR threshold used here.

